



Planck is >20 Years of Work



2010-11-04 Ken Ganga

(A)

Telescope

Primary: 1.50 x 1.89 m ellipsoid (CFRP)

Secondary: 1.02 x 1.04 m ellipsoid (CFRP)

System:

1.5 m circular projected aperture

Total MWFE<40 μm ms

- Total ε <0.01

 Reflectors are being developed by Astrium Gmbh, led by ESA and the Danish Space

Research Institute

(PI: Dr. H.U. Norgaard-Nielsen)



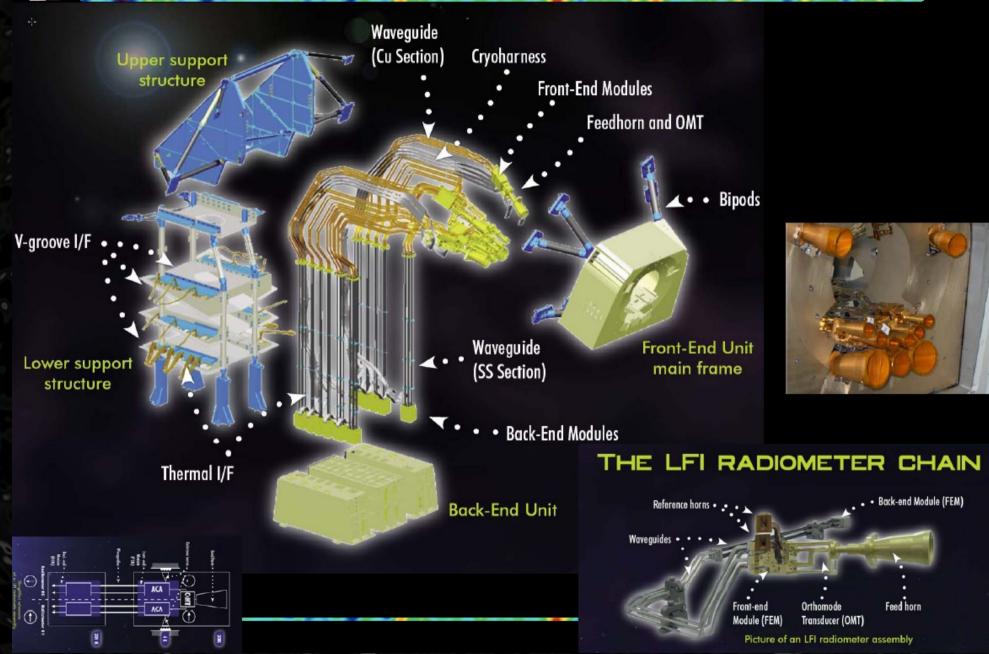






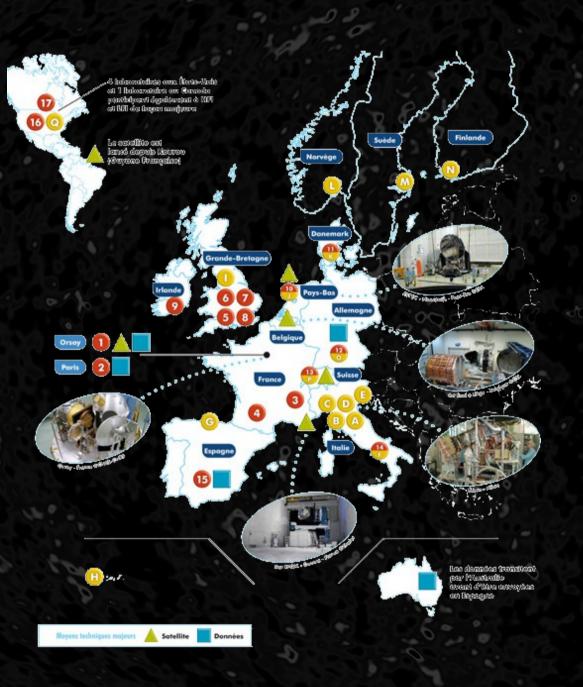
The Low Frequency Instrument LFI







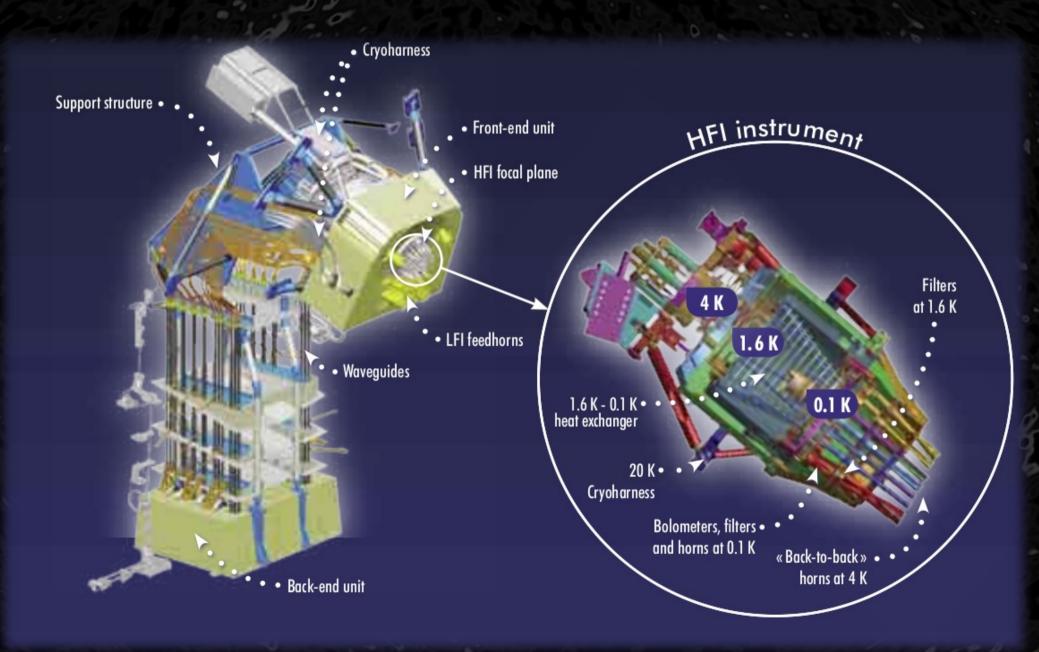
Contributions to Planck and HFI





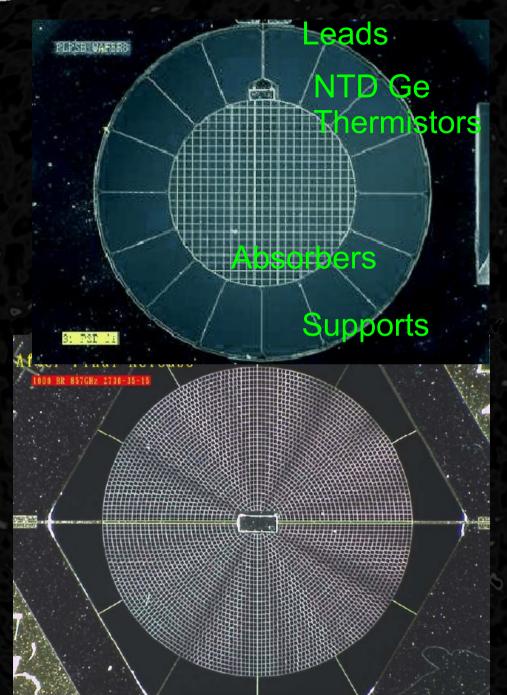


The Coldest Thing in Space





Polarization and Total Power Detect.

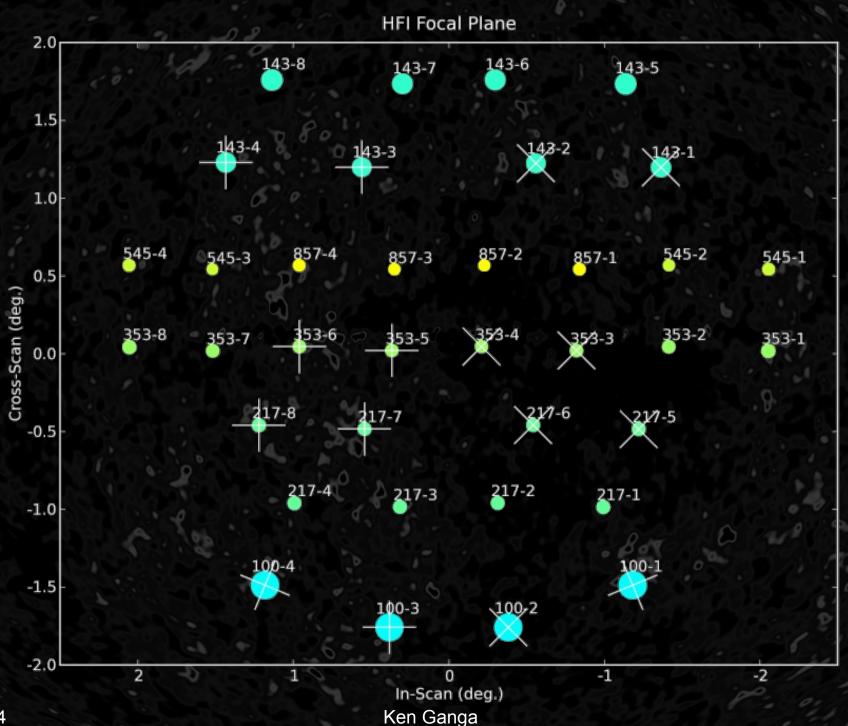


There are actually two detectors in the top figure to the left – each sensitive to a single linear polarization

Effort is required to understand the "cross-polarization" of the detectors



The HFI Focal Plane: ~50 Detectors

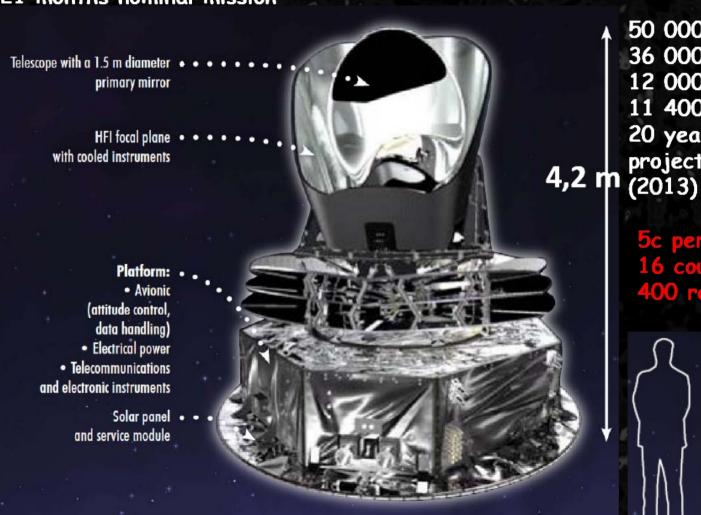




Payload Summary

2000 Kg 1600 W consumption 2 instruments - HFI & LFI 21 months nominal mission





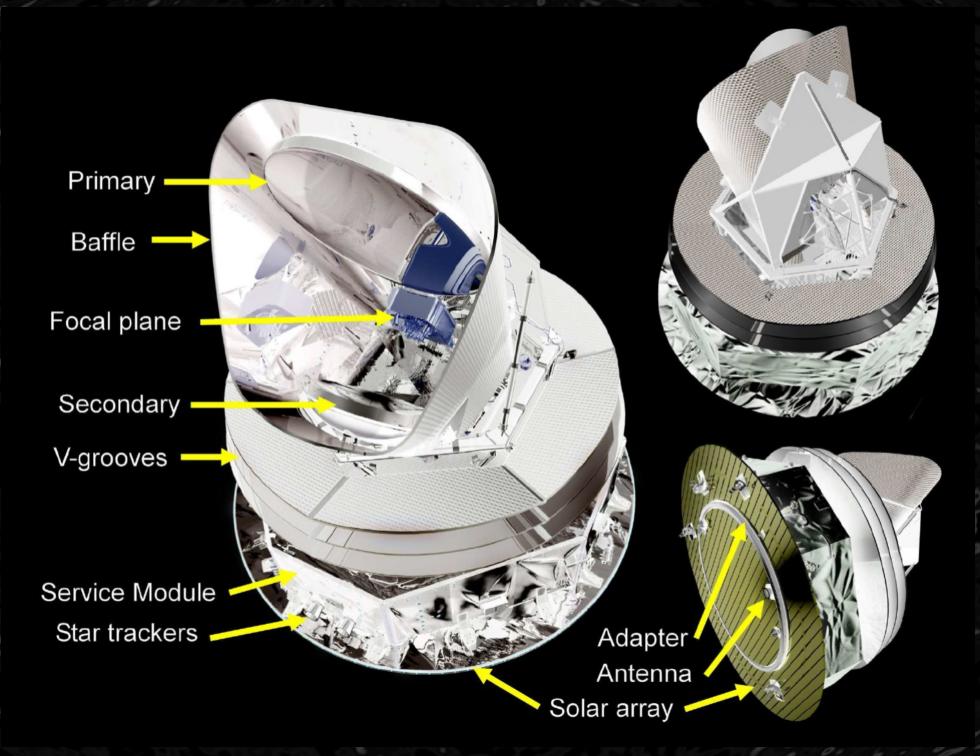
50 000 electronic components 36 000 1 4He 12 000 l 3He 11 400 documents 20 years between the first project and first results

per European per year 16 countries 400 researchers among 1000



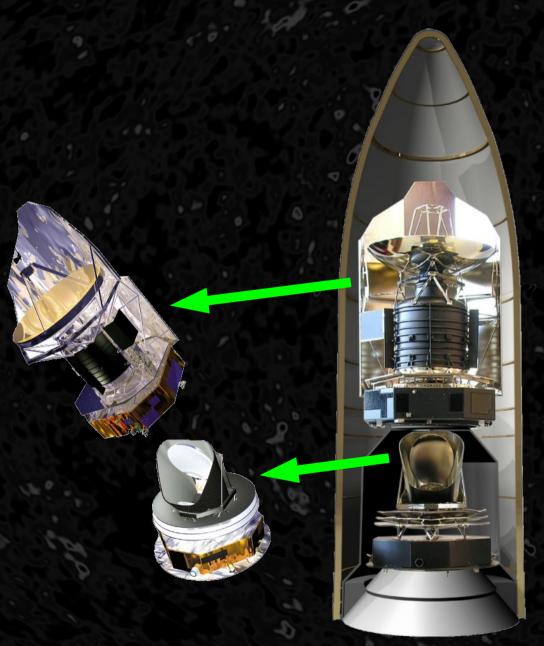
4,2 m Ken Ganga







Launch

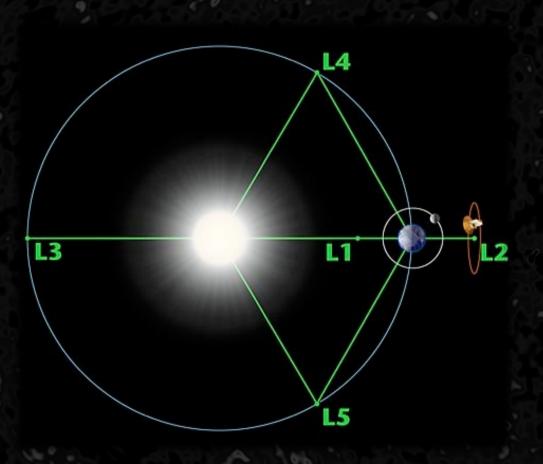


Launch was May 1 from Guiana Space Center, Kourou, French Guiana Ariane 5 Launch Vehicle With Herschel, which also is at L-2





2nd Earth-Sun Lagrange Point



- Radiation from the Earth and the atmosphere can contaminate our measurements
- Solution: Get as far from the Earth as possible
- L2 is ~1.5·10⁶ km (~0.01 AU) from Earth

http://map.gsfc.nasa.gov



Planck's Orbit and Scanning

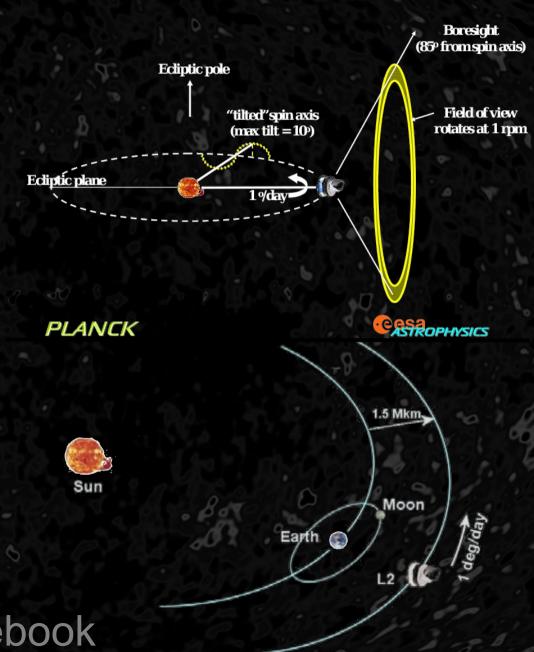
Orbits the Sun

Solar Panels always facing near Sun

Antenna always facing near Earth

Detectors always ~95° from Sun/Earth

We actually orbit around L2 (1.5 Mkm from Earth) in a Lissajous orbit



Planck Bluebook

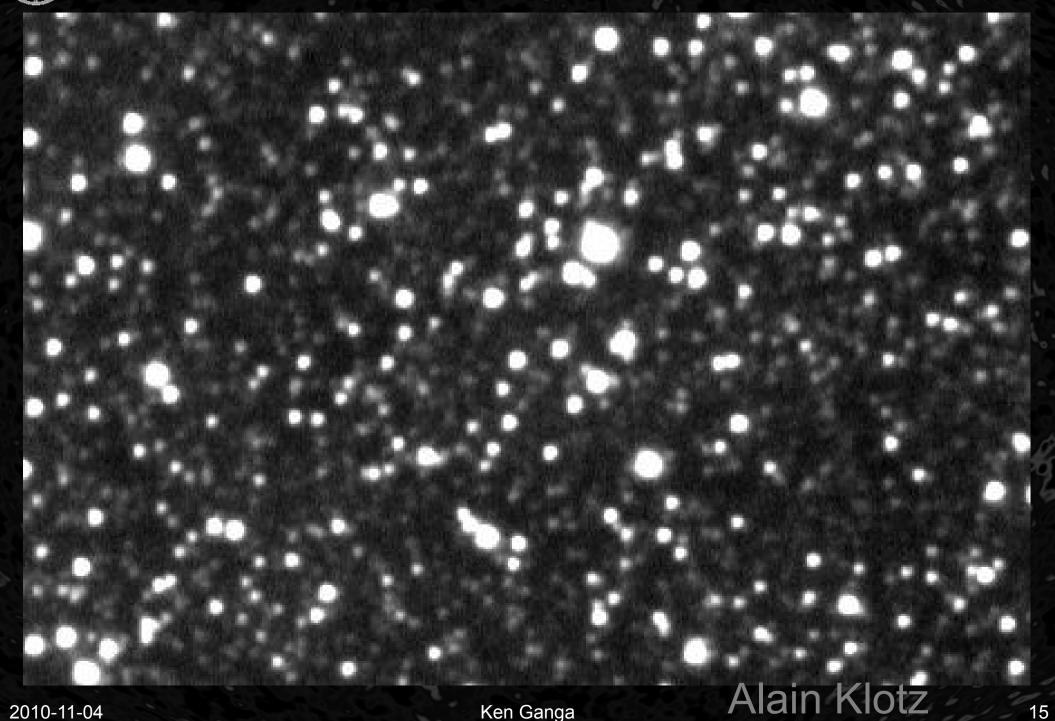


Voyage to Orbit

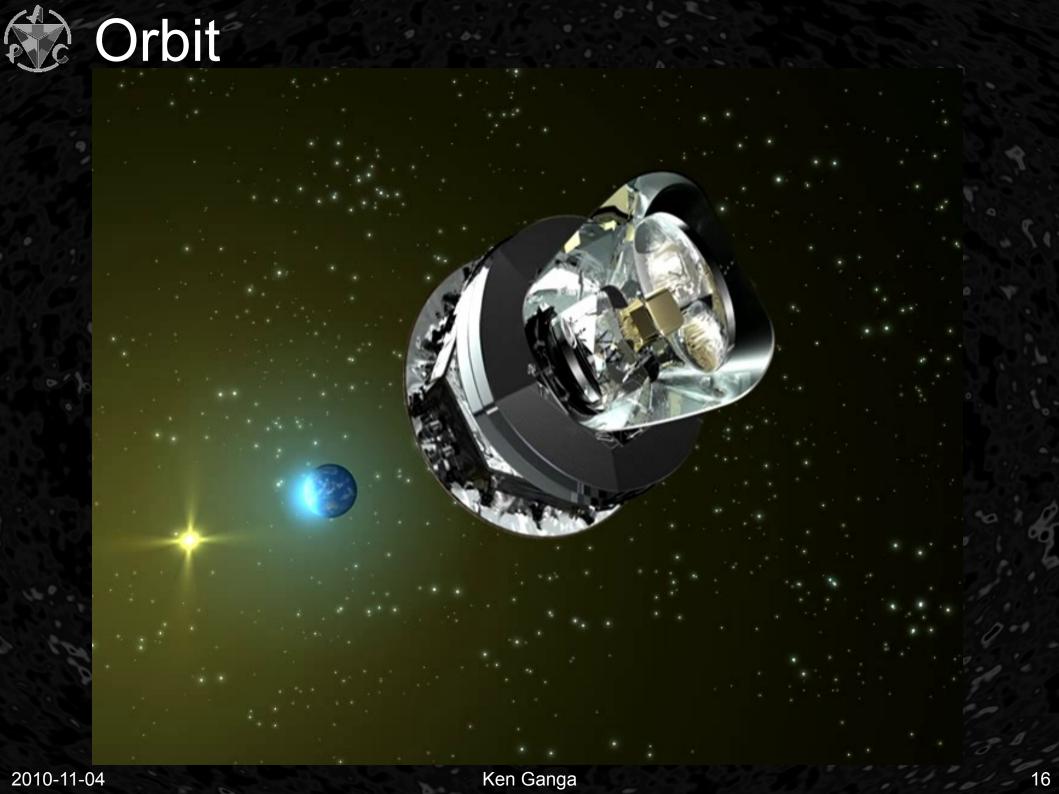




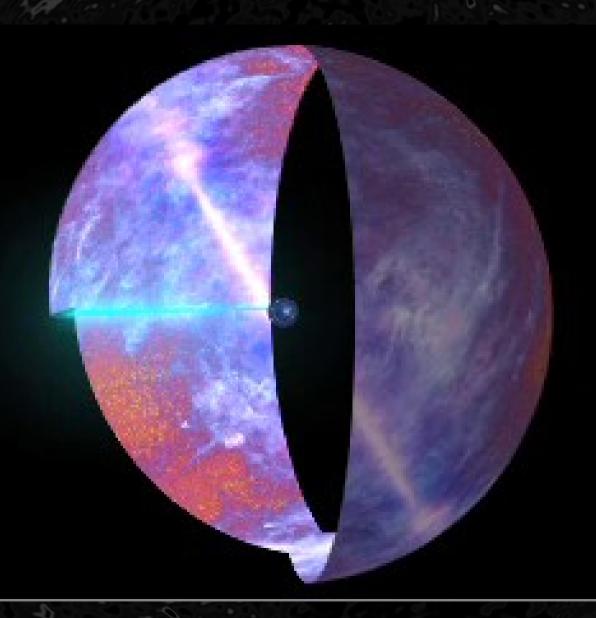
Planck as Seen from Earth



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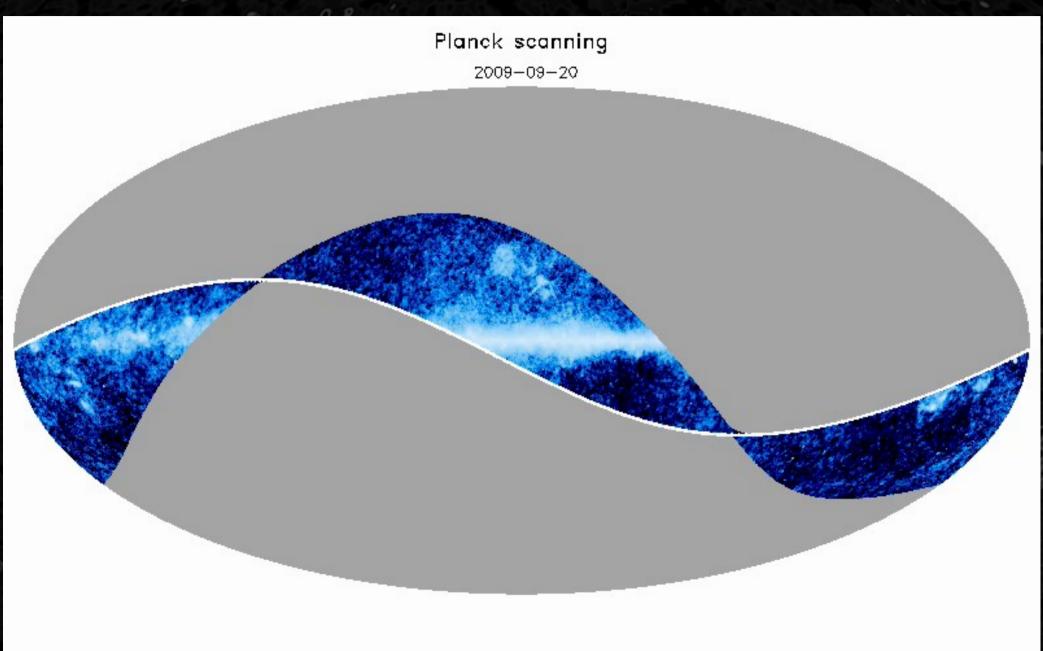




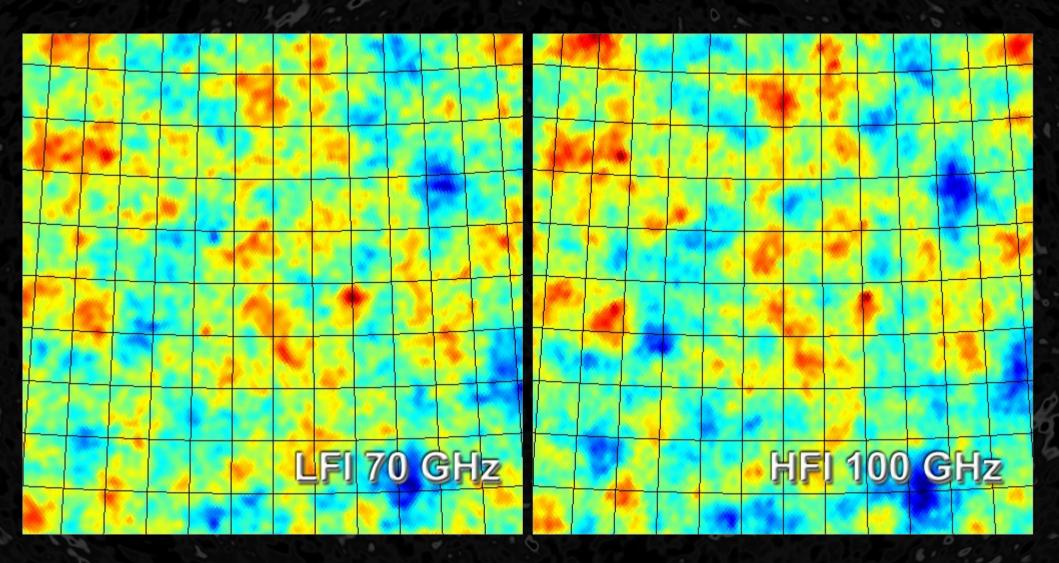




Sky Coverage

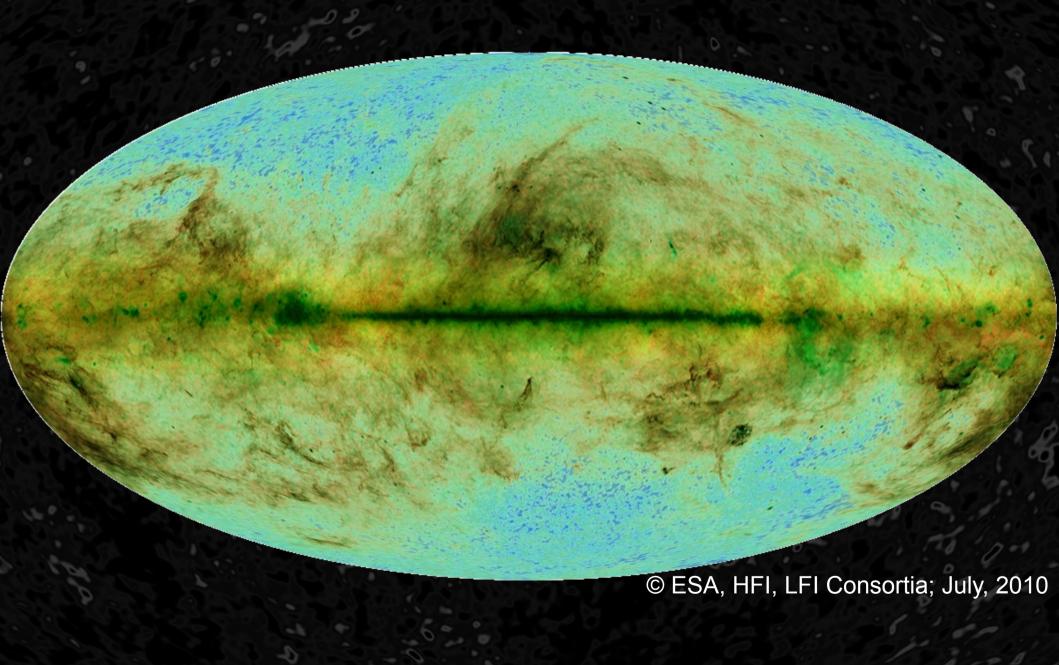






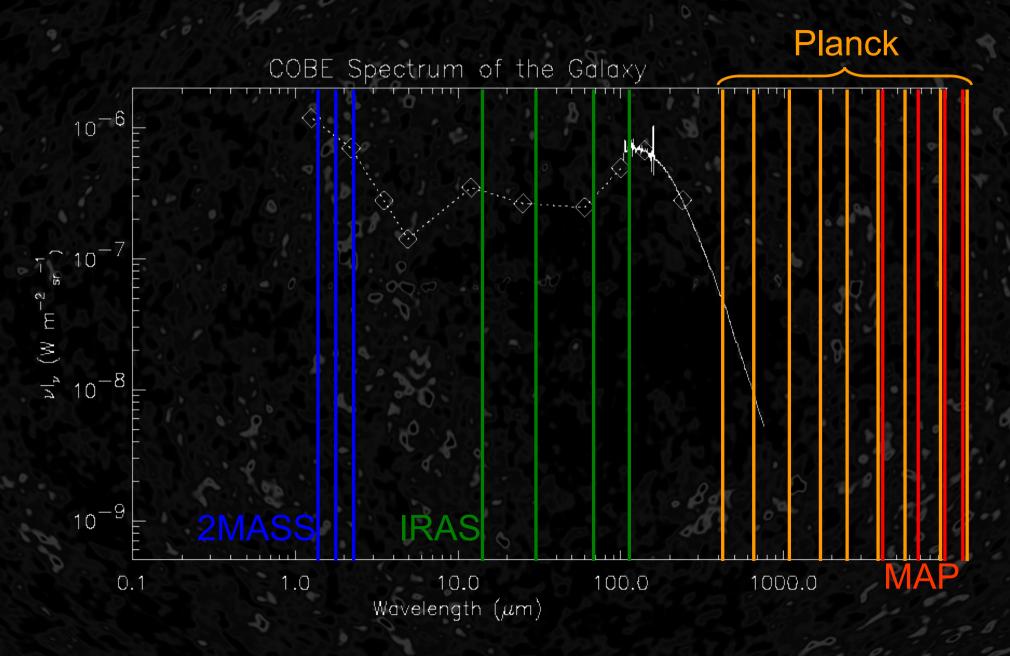


Planck One-Year All-Sky Survey





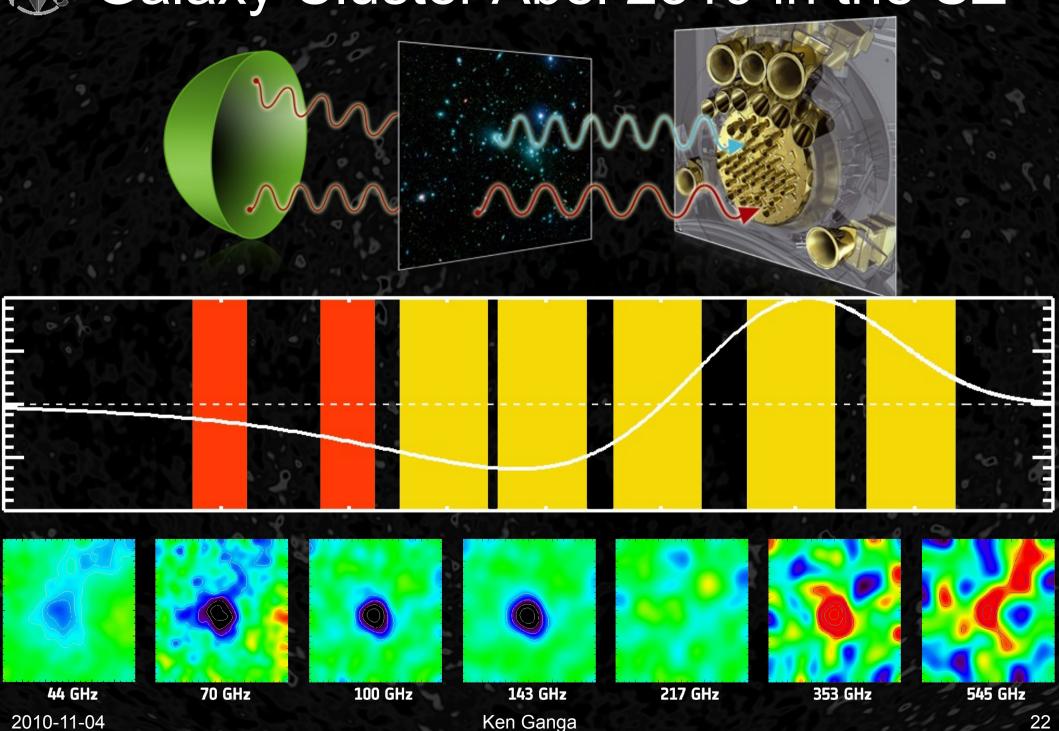
Planck Frequency Coverage



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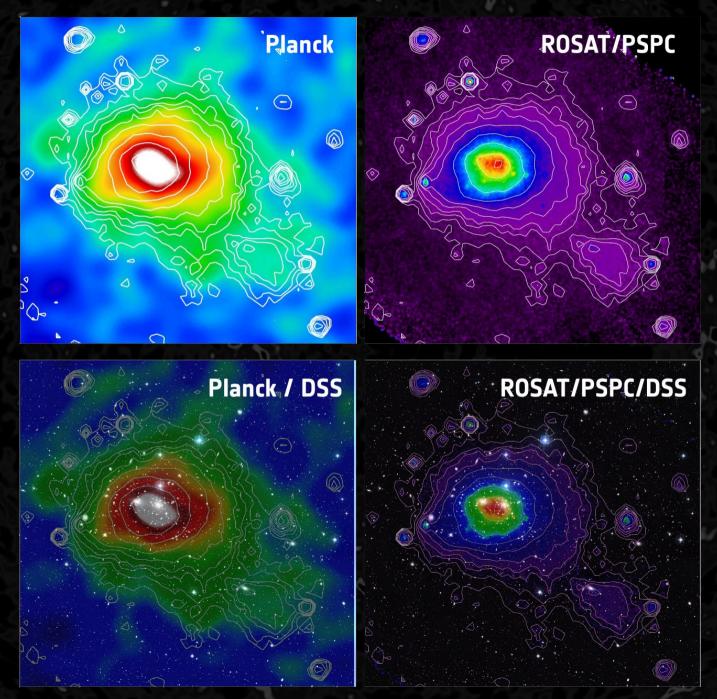


Galaxy Cluster Abel 2319 in the SZ





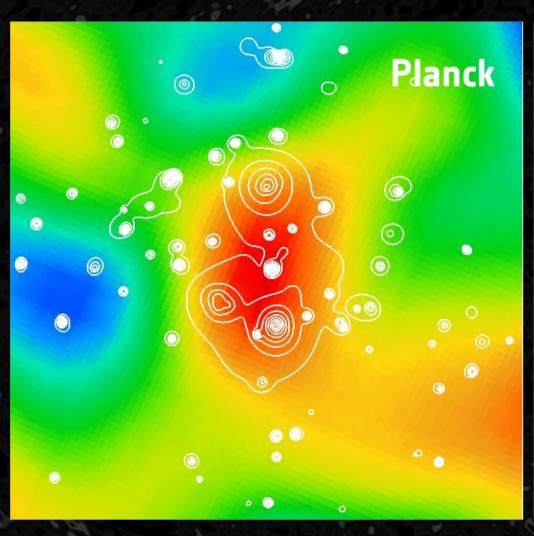
The Coma Cluster



23



A New Supercluster Seen By Planck



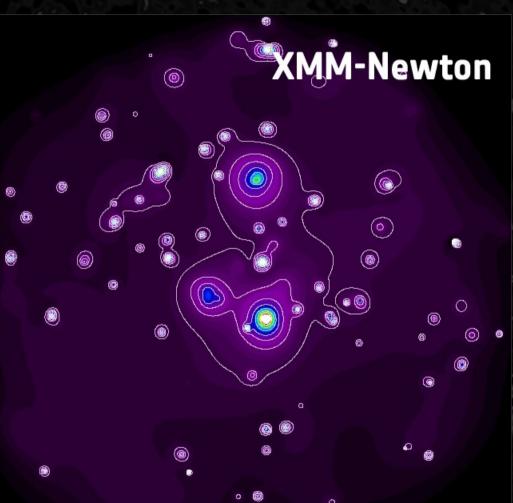




Illustration of polarization sensitivity

WMAP 7 years (Komatsu et al. preprint 2010) & simulations

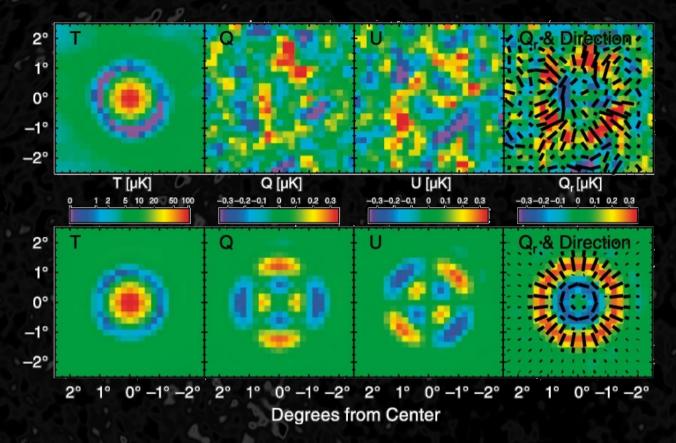
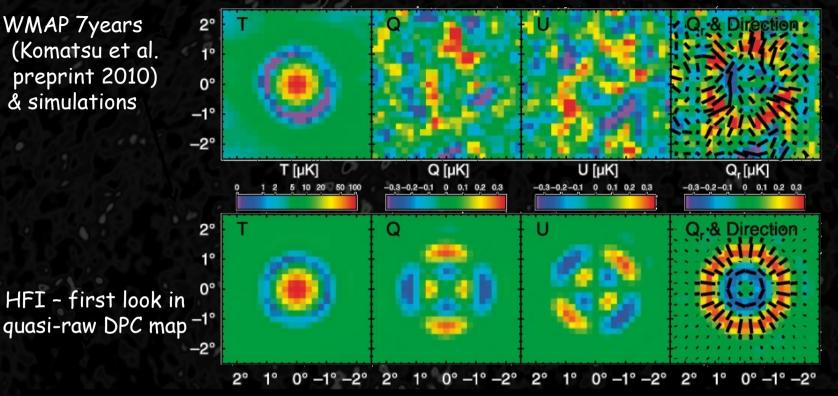




Illustration of polarization sensitivity





Health Warning: not assured yet whether systematic effects are controlled for precision cosmology

ILLUSTRATIVE Planck PLOTS REMOVED FOR DISTRIBUTION



WMAP/Planck E-Mode Polarization

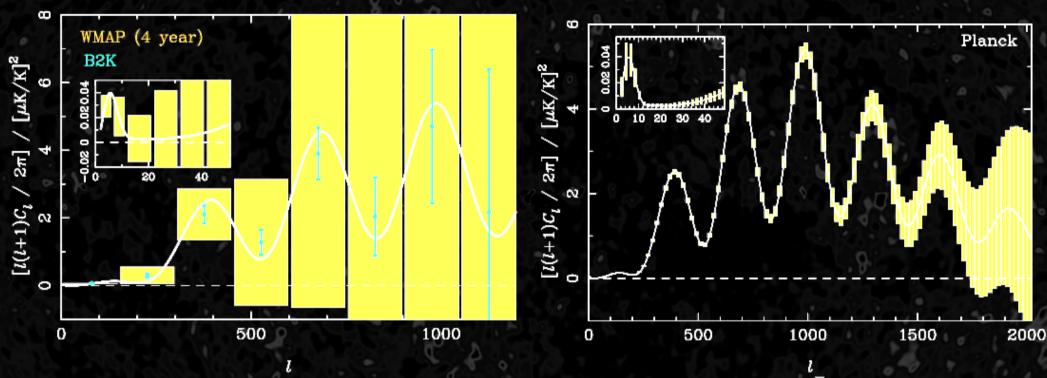


FIG 2.14.—Forecasts for the $\pm 1\sigma$ errors on the *E*-mode polarization power spectrum C_{ℓ}^{E} from *WMAP* and B2K (left) and *Planck* (right). The cosmological model, and the assumptions about instrument characteristics, are the same as in Figure 2.13. For *WMAP* and B2K, flat band powers are estimated with $\Delta \ell = 150$ (with finer resolution on large scales for *WMAP* in the inset). For *Planck* we have used the same ℓ -resolution as in Figure 2.13.



Projected Tensor-to-Scaler Limits

B-mode Detection with an Extended Planck Mission

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Kavli Institute for Cosmology Cambridge and Institute of Astronomy, Madingley Road, Cambridge, CB3 OHA.

Abstract. The *Planck* satellite has a nominal mission lifetime of 14 months allowing two complete surveys of the sky. Here we investigate the potential of an extended *Planck* mission of four sky surveys to constrain primordial B-mode anisotropies in the presence of dominant Galactic polarized foreground emission. An extended Planck mission is capable of powerful constraints on primordial B-modes at low multipoles, which cannot be probed by ground based or sub-orbital experiments. A tensor-scalar ratio of r = 0.05 can be detected at a high significance level by an extended Planck mission and it should be possible to set a 95\% upper limit of $r \lesssim 0.03$ if the tensorscalar ratio is vanishingly small. Furthermore, extending the Planck mission to four sky surveys offers better control of polarized Galactic dust emission, since the 217 GHz frequency band can be used as an effective dust template in addition to the 353 GHz channel.



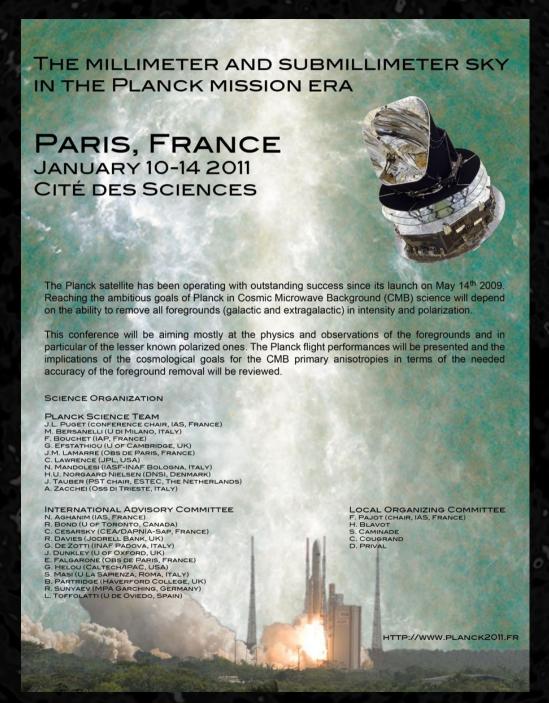
Schedule (unofficial)



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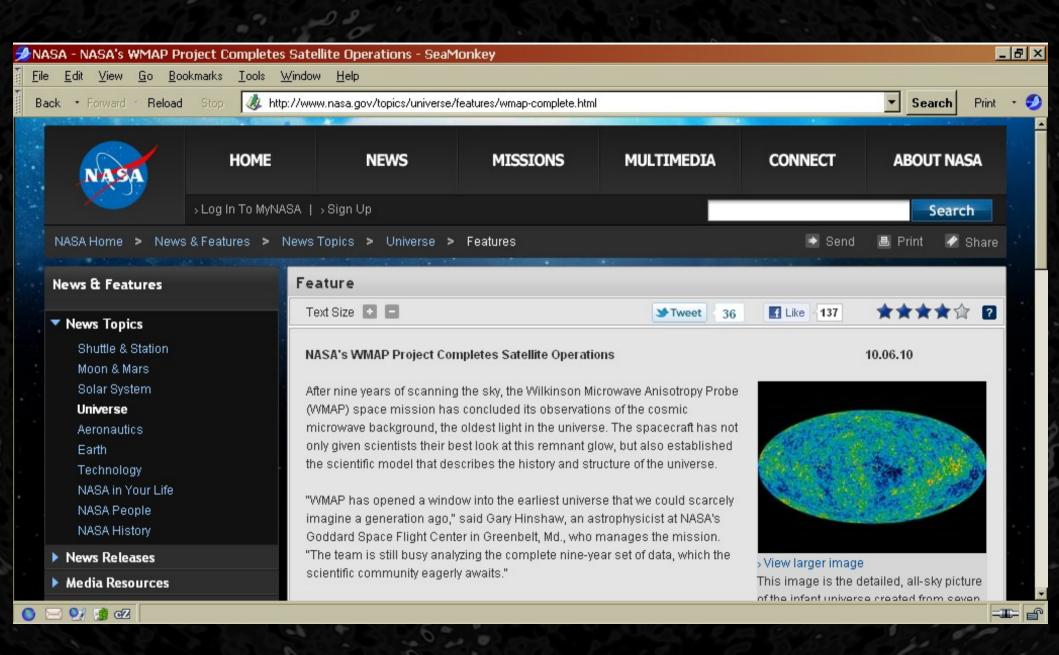


A First Planck Conference





Planck Continues as WMAP Retires







Thank You!

